

A Novel Update Algorithm in Stokes Space for Adaptive Equalization in Coherent Receivers

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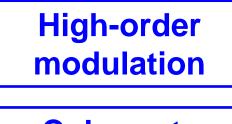
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Coherent detection and DSP



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Coherent detection

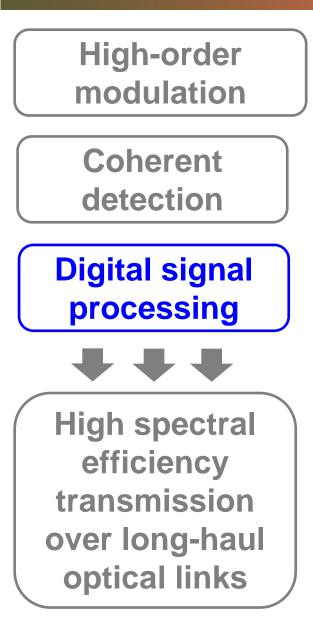
Digital signal processing



High spectral efficiency transmission over long-haul optical links



Coherent detection and DSP

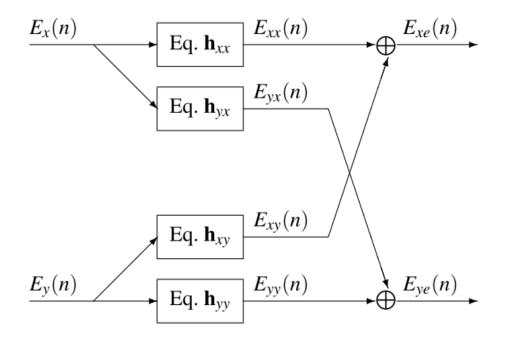


- Front-end IQ imbalance
- Chromatic dispersion
- PMD
- PDL
- Residual CD
- Filtering effects
- Frequency offset
- Laser phase noise
- Fiber non-linear effects

Adaptive MIMO equalizer



MIMO equalizer



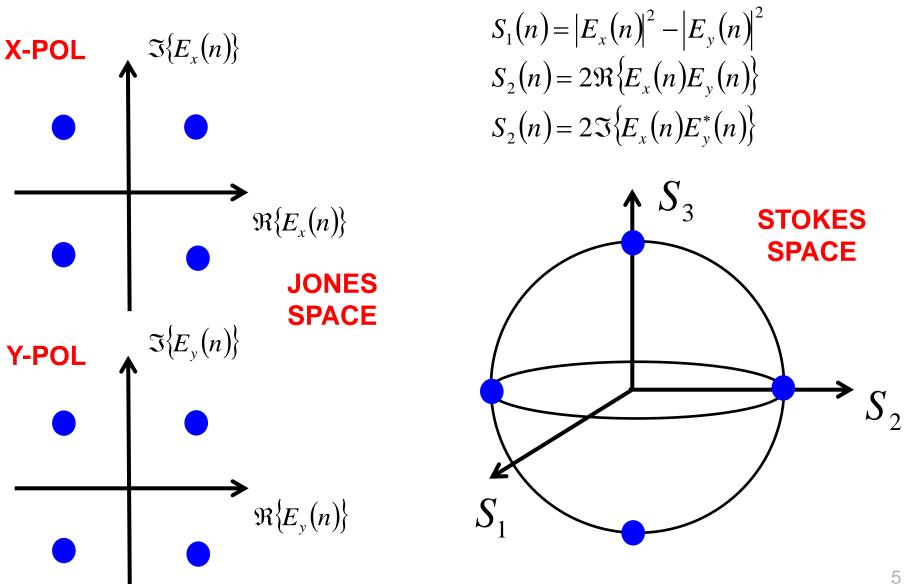
- $E_{xe}(n) = \mathbf{E}_{x}^{T}\mathbf{h}_{xx} + \mathbf{E}_{y}^{T}\mathbf{h}_{xy}$ $E_{ye}(n) = \mathbf{E}_{x}^{T}\mathbf{h}_{yx} + \mathbf{E}_{y}^{T}\mathbf{h}_{yy}$
- Standard CMA or LMS algorithms: update of the coefficients based on error signals evaluated on the two-dimensional constellations (separate for the two polarizations)

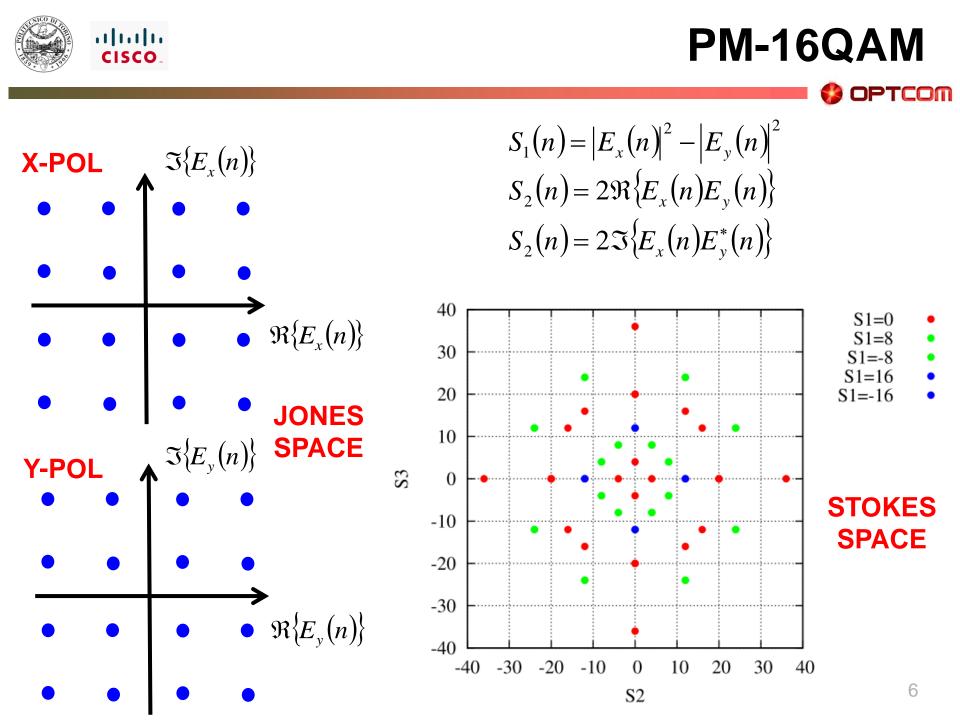
New algorithm: update of the coefficients based on error signal evaluated in the Stokes space

Performance test on a PM-16QAM signal, comparing it to the multi-modulus CMA algorithm



PM-QPSK







Design rule Stokes equalizer

Error function to be minimized

$$f(\mathbf{h}) = f(\mathbf{h}_{xx}, \mathbf{h}_{xy}, \mathbf{h}_{yx}, \mathbf{h}_{yy}) = \\ = \left(S_{1e}(n) - \hat{S}_{1}(n)\right)^{2} + \left(S_{2e}(n) - \hat{S}_{2}(n)\right)^{2} + \left(S_{3e}(n) - \hat{S}_{3}(n)\right)^{2}$$

with:

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$$\mathbf{S}_{e} = [S_{1e}(n), S_{2e}(n), S_{3e}(n)]$$

$$\hat{\mathbf{S}} = \left[\hat{S}_1(n), \hat{S}_2(n), \hat{S}_3(n)\right]$$

Stokes vector of the equalized signal

Stokes vector of the transmitted signal

either known (training sequence) or estimated (decisiondirected)

PTCOM

111111 Taps update algorithm – Stokes/CMA cisco

Rule for adaptively update the equalizer weights:

 $\mathbf{h}_{xx}(n+1) = \mathbf{h}_{xx}(n) - \mu \nabla_{\mathbf{h}} f(\mathbf{h}(n))$ $\mathbf{h}_{xv}(n+1) = \mathbf{h}_{xv}(n) - \mu \nabla_{\mathbf{h}_{xv}} f(\mathbf{h}(n))$ $\mathbf{h}_{vx}(n+1) = \mathbf{h}_{vx}(n) - \mu \nabla_{\mathbf{h}_{vy}} f(\mathbf{h}(n))$ $\mathbf{h}_{vv}(n+1) = \mathbf{h}_{vv}(n) - \mu \nabla_{\mathbf{h}_{m}} f(\mathbf{h}(n))$

Stokes algorithm

$$\begin{bmatrix} C_1(n) \\ C_2(n) \end{bmatrix} = \begin{bmatrix} \varepsilon_1(n) & \varepsilon_2(n) \\ \varepsilon_2^*(n) & -\varepsilon_1(n) \end{bmatrix} \begin{bmatrix} E_{xe}(n) \\ E_{ye}(n) \end{bmatrix}$$
$$\varepsilon_1(n) = S_{1e}(n) - \hat{S}_1(n)$$
$$\varepsilon_2(n) = \left(S_{2e}(n) - \hat{S}_2(n)\right) + j\left(S_{3e}(n) - \hat{S}_3(n)\right)$$

Evaluation of gradients:

$$\nabla_{\mathbf{h}_{xx}} f(\mathbf{h}(n)) = C_1(n) \mathbf{E}_x^*$$
$$\nabla_{\mathbf{h}_{xy}} f(\mathbf{h}(n)) = C_1(n) \mathbf{E}_y^*$$
$$\nabla_{\mathbf{h}_{yy}} f(\mathbf{h}(n)) = C_2(n) \mathbf{E}_y^*$$
$$\nabla_{\mathbf{h}_{yx}} f(\mathbf{h}(n)) = C_2(n) \mathbf{E}_x^*$$

• CMA

$$\begin{bmatrix} C_1(n) \\ C_2(n) \end{bmatrix} = \begin{bmatrix} \varepsilon_x(n) & 0 \\ 0 & \varepsilon_y(n) \end{bmatrix} \begin{bmatrix} E_{xe}(n) \\ E_{ye}(n) \end{bmatrix}$$

$$\varepsilon_x(n) = |E_{xe}(n)|^2 - R^2$$

$$\varepsilon_y(n) = |E_{ye}(n)|^2 - R^2$$



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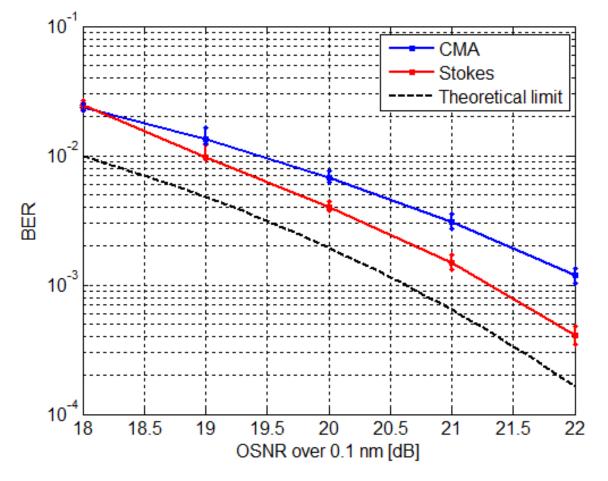
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Case study – PM-16QAM

- Symbol rate: R_s=32 Gbaud
- Single-channel
- Nyquist spectrum (raised-cosine with roll-of 0.1)
- Residual CD = 250 ps/nm
- DGD = 1 symbol
- BER values estimated through Monte-Carlo simulation for several combinations of DGD axis and state of polarization (SOP) at the input of the Rx, for a total of ~900 cases
- Equalization using a training sequence, followed by decisiondirected operation



BER vs. OSNR

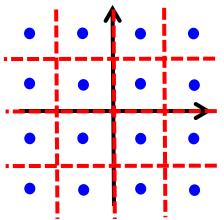


- Number of eq. taps: M = 31.
- Solid lines: average performance
- Vertical bars: range of variations for all considered values of SOP and DGD axis.
- The value of the adaptive equalizer update coefficient μ was optimized for both CMA and Stokes algorithms.



Performance improvement

- Performance can be improved by:
- 1. Using an adaptive Maximum-Likelihood decision criterion instead of a fixed-threshold one



2. Changing the decision rule in the Stokes space (minimum distance is not optimum)

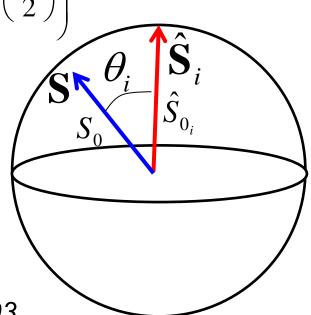
Statistics of noise in Stokes space

•
$$\mathbf{S} = (S_1, S_2, S_3)$$
 = noisy received vector
• $\hat{\mathbf{S}}_i = (\hat{S}_{1_i}, \hat{S}_{2_i}, \hat{S}_{3_i})$ = ideal un-noisy constellation vecto

$$\blacktriangleright \mathsf{PDF} \text{ of } \mathbf{S} \Big| \widehat{\mathbf{S}}_{i} [*]: f_{\mathbf{S}|\widehat{\mathbf{S}}_{i}} = \frac{e^{-\frac{S_{0i} + S_{0}}{2\sigma^{2}}}}{16\pi S_{0}\sigma^{4}} I_{0} \left(\frac{\sqrt{\widehat{S}_{0i} \cdot S_{0}}}{\sigma^{2}} \cos\left(\frac{\theta_{i}}{2}\right)\right)$$

- $\bullet S_0$ = magnitude of S
- \hat{S}_{0_i} = magnitude of \hat{S}_i
- \bullet θ_i = angle between **S** and $\hat{\mathbf{S}}_i$
- σ^2 = noise variance in each polarization

[*] P. Poggiolini, "Digital optical transmission systems based on polarization modulation", PhD thesis, 1993





Decision rule



$$f_{\mathbf{S}|\widehat{\mathbf{S}}_{i}} = \frac{e^{-\frac{\widehat{S}_{0_{i}} + S_{0}}{2\sigma^{2}}}}{16\pi S_{0}\sigma^{4}} I_{0} \left(\frac{\sqrt{\widehat{S}_{0_{i}} \cdot S_{0}}}{\sigma^{2}} \cos\left(\frac{\theta_{i}}{2}\right)\right)$$

- $\mathbf{S} = \begin{pmatrix} \mathbf{\hat{S}}_{i} & \mathbf{\hat{S}}_{i} \\ \mathbf{\hat{S}}_{0} & \mathbf{\hat{S}}_{0i} \end{pmatrix}$
- The decision rule can be based on the maximization of $f_{\mathbf{S}|\widehat{\mathbf{S}}_i}$ over all possible noiseless constellation points $\widehat{\mathbf{S}}_i$.
- There are common factors across all possible indices i that can be eliminated → we can apply the ML decision on the formula:

$$p_i = e^{-\frac{S_0}{2\sigma^2}} I_0 \left(\frac{\sqrt{\hat{S}_{0_i} \cdot S_0}}{\sigma^2} \cos\left(\frac{\theta_i}{2}\right) \right)$$



Taking the logarithm and applying some simplifications, we obtain the following new metric (based on actual statistics in Stokes space):

$$m_i \approx -S_{0_i} + 2\sqrt{\hat{S}_{0_i}}\sqrt{S_0} \cos\left(\frac{\theta_i}{2}\right)$$
 NOVEL METRIC

 Minimum-distance metric (based on Gaussian distribution hypothesis)

$$d_{i}^{2} = \left(S_{1} - \hat{S}_{1_{i}}\right)^{2} + \left(S_{2} - \hat{S}_{2_{i}}\right)^{2} + \left(S_{3} - \hat{S}_{3_{i}}\right)^{2} = S_{0}^{2} + \hat{S}_{0_{i}}^{2} - 2\hat{S}_{0_{i}}S_{0}\cos(\theta_{i})$$
$$\implies -\hat{S}_{0_{i}}^{2} + 2\hat{S}_{0_{i}}S_{0}\cos(\theta_{i}) \qquad \text{MINIMUM DISTANCE} \\ \text{METRIC}$$

16QAM with adaptive ML receiver

- 10^{-1} CMA Stokes - new metrics Theoretical limit 10⁻² 10⁻³ 4 10 17 18 19 20 21 16 OSNR over 0.1 nm [dB]
- Number of eq. taps: M = 31.

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 Solid lines: average performance (over all considered values of SOP and DGD axis).

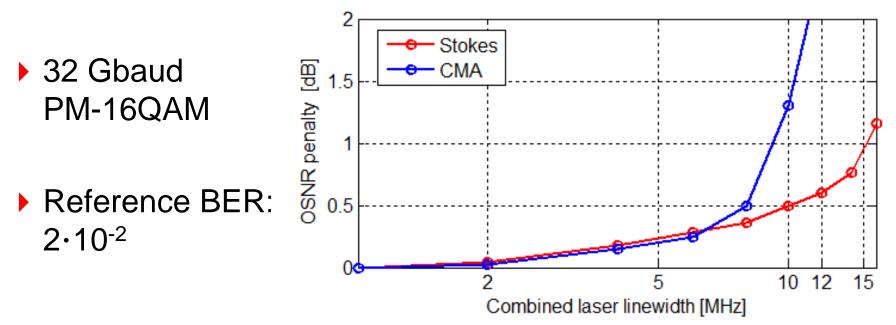
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Phase noise tolerance

- CPE block (inserted after the MIMO equalizer), based on the Viterbi&Viterbi algorithm, with QPSK partitioning
- Stokes-space update guarantees that the two polarizations after equalization are perfectly aligned to each other in phase → the phase error estimate can be obtained as an average on both polarizations





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Conclusions

- A novel update algorithm for adaptive MIMO equalizer taps has been proposed and its performance analyzed in a PM-16QAM scenario.
- At the expenses of a slight increase in complexity, it has the following advantages with respect to standard CMA and LMS algorithms:

With respect to LMS

Insensitive to phase noise and frequency offset

With respect to CMA

- Not affected by the "degeneracy" problem typical of CMA
- ► CPE algorithms can estimate the phase noise by averaging over the two polarizations → nearly doubled phase-noise tolerance







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